

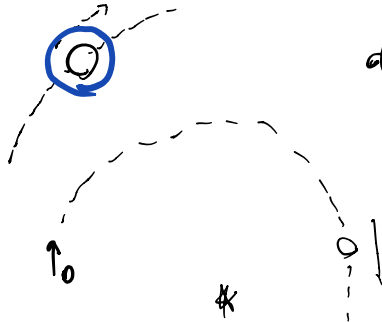
Velocità della luce finita



Romer e la luna di Giove "Io"

$$v_{\oplus} \sim 10^{-2} c$$

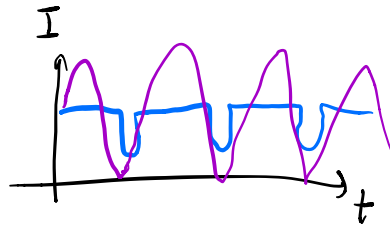
$$v_{\odot} \sim 10^{-3} c$$



altri sistemi di satelliti
danno la stessa c
 $\Rightarrow c \ll \infty$ e non
dipende dalla vel. rel.
tra sorgente e osservatore

$$\tau_{zw} = 3 - 15 \text{ min} \quad \tau_{zw} = 3 + 15 \text{ min}$$

Stelle binarie

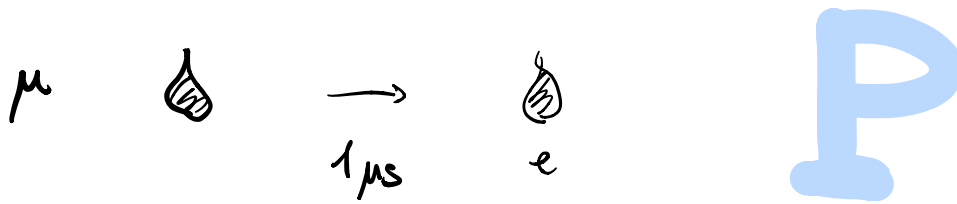


c non dipende
da v_{sorgente}

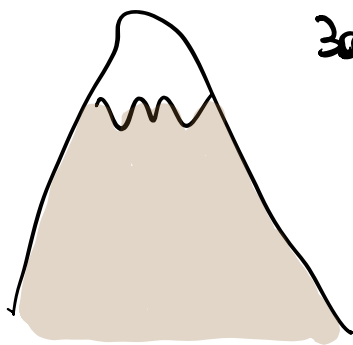
Michelson & Morely

$$c \cong 30 \text{ cm/ms}$$

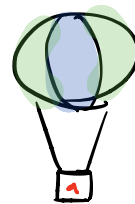




$$d = 1000 \cdot 20 \text{ cm} = 300 \text{ m} \left(\frac{v}{c} \right)$$



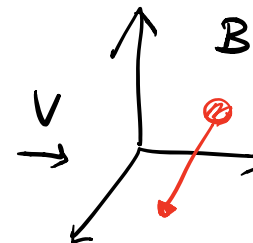
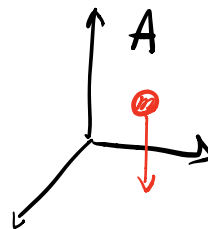
3000 m — N_{3000}



$\sim 150 - 200 \text{ m} - N_0 \approx N_{2000}$

$v = c$ è un fatto assoluto

$$v'_B = v_A + V$$



$$x' = [x + vt] \gamma(v)$$

↓

$$\Delta x' = (\Delta x + v \Delta t) \gamma(v)$$

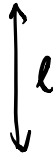
$$\omega' = \frac{\Delta x'}{\Delta t'} = \left(\frac{\Delta x}{\Delta t} + v \frac{\Delta t}{\Delta t'} \right) \gamma(v)$$

$$\Delta t = \Delta t' \quad (\text{temps universel})$$

$$\omega' = (\omega + v) \gamma(v)$$

$$\gamma(v) = \frac{1}{1 + v/c}$$

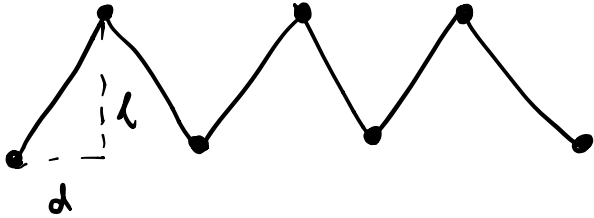
ARRETRANDO A FORMA



$$v = 0$$

$$t_0 = 2l/c$$

$$l = ct_0$$



$$D = \sqrt{l^2 + d^2} \quad t = D/c$$

$$t^2 = \frac{l^2 + d^2}{c^2} = t_0^2 + \frac{d^2}{c^2} = t_0^2 + \left(\frac{v}{c}t\right)^2$$

$$t_0^2 = t^2 - \frac{v^2}{c^2} t^2$$

$$t_0 = t(1 - \beta^2)^{1/2} = t/\gamma$$

$$\Delta t' = \Delta t_0 (1 - \beta^2)^{-1/2}$$

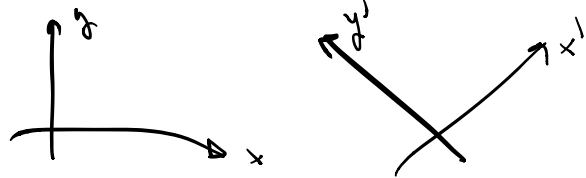
$$t' = \left(t - \beta \frac{x}{c}\right) \gamma$$

IL TEMPO NEW È ASSOLUTO

$$[x] = [\beta \cdot t c] \quad \beta = \frac{v}{c} \ll 1$$

$$[t] = \left[\beta \frac{x}{c} \right]$$

$$\left\{ \begin{array}{l} x' = \frac{x - \beta t c}{\sqrt{1 - \beta^2}} \approx (x - \beta t c) \left(1 + \frac{1}{2} \beta^2 + \dots \right) \\ t' = \frac{t - \beta \frac{x}{c}}{\sqrt{1 - \beta^2}} \end{array} \right.$$



$$x^2 + y^2 = x'^2 + y'^2$$

$$c^2 t'^2 - x'^2 = c^2 \frac{t^2 - 2\beta \frac{x}{c} t + \beta^2 \frac{x^2}{c^2}}{1 - \beta^2} - \frac{x^2 - 2\beta x t c + \beta^2 x^2 c^2}{1 - \beta^2}$$

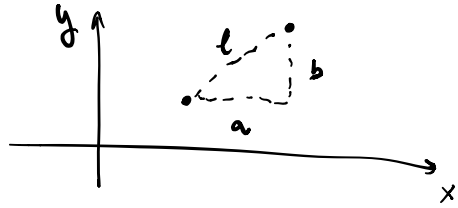
$$= \frac{c^2 t^2 (1 - \beta^2) - x^2 (1 - \beta^2) - 2\beta c (x t - x t)}{1 - \beta^2}$$

$$= c^2 t^2 - x^2$$

$$c^2 t'^2 - x'^2 = c^2 t^2 - x^2$$

$$l = \sqrt{a^2 + b^2}$$

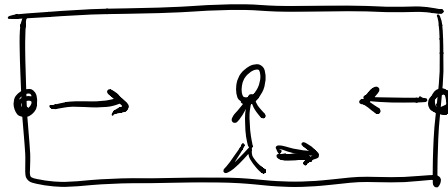
$$s^2 = a^2 - b^2$$



$$s^2 = t^2 - x^2 = t'^2 - x'^2 = s'^2$$

↑
intervalli

SIMULTANEITA'



$$\Delta t^{\text{trav}} = 0$$

$\Delta t \geq 0$ dalla formula

$$x^\mu = (t, \vec{x}) = \Lambda x^\mu$$

$$\partial_\mu = \frac{d}{dx^\mu} \sim \Lambda^{-1}$$

$$i\partial_\mu = P_\mu = (E, \vec{p}_x)$$

contrazione dei tempi \rightarrow aumento delle energie

contrazione della lunghezza \rightarrow aumento delle q.ta di moto



$$\gamma m \cdot v \left(\frac{1}{\gamma} \right) = p$$



perché $mv = mv$ in ciascun sistema di riferimento

in cui $m \neq m$, v dell'altro deve variare minore

a causa della dilatazione dei tempi, $\frac{1}{\gamma} v' = v/\gamma$

Se $m \cdot v = m' \cdot v'$ per conservare la q.ta di

moto allora m deve aumentare in $m' = m\gamma$

$$m\gamma = m \frac{1}{\sqrt{1-\beta^2}} = m_0 \left(1 + \frac{1}{2}\beta^2\right)$$

$$m c^2 = m_0 c^2 + \underbrace{\frac{1}{2} m_0 v^2}_{E_{\text{kin}}}$$

$$E_{\text{tot}} = E_{\text{rest}} + E_{\text{kin}} = m_0 \gamma c^2$$